the two stations indicates a degree of relationship so slight as to be negligible. During this period the range in the monthly mean values was from 1.933 to 1.907, or

0.026 gr. cal.
On Figures 2 and 3 a cross has been located within the circles that inclose most of the dots, to show the mean value of the solar constant determinations indicated by the dots. A second cross has been located at a point to indicate the mean of solar constant values represented by the dots on the other corresponding figure. These two crosses fall on a line making an angle of very nearly 45° with the axes of ordinates and abscissas, which indicates that the results as summarized above are not inconsistent with Abbot's finding of a high correlation coefficient between the means of groups of synchronous values at Montezuma and Harqua Hala, arranged in order of magnitude of the solar constant determinations at Montezuma.8 Such a grouping eliminates very largely the day-to-day variations discussed above; and the correlations found are the result of secular changes occurring during the period covered by the three groups, and not real correlations between actual dayto-day changes free from secular variations.

Furthermore, during the period April, 1922, to November, 1924, inclusive, a solar constant value is given on 827 days, and on 299 of these days, or on 1 day out of 2.78, values obtained at both stations were included in the mean for the day. The mean value for the whole period is 1.922. On 36 days the value was less than 1.900, and on 6 of these days, or on only 1 day out of 6, was the value derived from determinations made at

⁸ Abbot, C. G., Solar variation and forecasting, p. 20, Smithsonian Misc. Coll., vol. 77, No. 5.

In view of the above, and for the further reason that one out of every six of these extreme values is by Abbot graded U+ or U (rather unsatisfactory or unsatisfactory), they are not entitled to as much weight as the

more nearly average values.

Abbot holds the view that since there are more values with large departures than theory calls for, this is a proof of solar variability. It is not unusual to find such an excess however. Further, in this particular case, it has been shown above that these extreme values have not the same degree of accuracy as the remaining values. Therefore the excess in their number, which is small numerically, can not be accepted as evidence of solar variability.

It seems evident, therefore, that the day-to-day variability of the solar constant determinations, the standard deviation of which is less than ± 0.70 per cent, depends largely upon whether the solar constant value is derived from determinations made at only one or at both stations; that it reflects unavoidable inaccuracies in pyrheliometric readings, and in extrapolating the readings to zero air mass; and that short-period solar variability, if it exists, falls within the limits of the probable error of the determinations.

THE PROBABLE 24-HOUR TEMPERATURE CHANGE (7 A. M. TO 7 A. M.) AT MONTGOMERY, ALA.

By Jesse W. Smith and Welby R. Stevens

[Montgomery, Ala., Weather Bureau Office, May 18, 1925]

In this study the probable temperature change for the 24-hour period 7 a. m. to 7 a. m., 90th meridian time, at Montgomery, Ala., has been determined by means of the Gram-Charlier frequency curves for each month of the year, based on 1,000 observations for each month.

The temperature changes were determined from the a. m. observations as recorded on Form No. 1001-Metl., beginning with 1924 and going back far enough to include 1,000 days in each monthly distribution. Each temperature was taken to the nearest even degree before the change was computed, thus giving the change in 2° units. This was done in order to give actual changes

considered in the verification of forecasts.

The Grain-Charlier curves were selected because of the relative ease with which they may be computed and their flexibility, which promised good fits in all cases. Reference to Figure 1 shows that very good fits were obtained. It is the belief of the authors that the Gram-Charlier curves are particularly well adapted to meteorological distributions, because of their capacity to take care of both skewness and excess, which are likely to be encountered in appreciable degree, especially in monthly distributions.

It seems unnecessary to give a detailed description of the method of fitting these curves. Reference is made to Fisher,² where a lucid explanation may be found, both

¹ Fisher, Arne, Mathematical Theory of Probabilities, New York, 1923.

tion. However, for practical application, tables will be necessary containing the values of the generating func-

tion
$$\varphi_0(z) = \frac{1}{\sqrt{2\pi}}e^{-z^2/2}$$
 and its derivatives up to at least

the fourth order.

Computations and the necessary control checks were made for each month, of which only January is shown, in Table 1, to illustrate the method.

Perhaps slightly better fits would have been obtained had the parameters been computed by the method of least squares instead of the method of semi-invariants, but the arithmetical labor entailed would have been

almost prohibitive.

It is well known that in dealing with a limited series of observations the third and fourth moments, or in this case the third and fourth semi-invariants, are liable to considerable error, due to the chance presence of a few large departures. It was found in most of the calculations that this error was sufficient to cause a slowing up in the rate of increase near the tails of the curves, or in some cases enough to cause serious secondary inflections. It was found possible to eliminate this undesirable situation by neglecting, in the computation of the third and fourth semi-invariants, the observations at the tails beyond the value $z = (x - \lambda_1) : \sigma > 4$. Inasmuch as never more

both stations. On 35 days the value was 1.940 or above, and on 4 of these, or on 1 day out of 8.75, the value was derived from determinations made at both stations. Thus, while more than one third of the daily values have been derived from measurements made at both stations, this is the case with only about one seventh of the values that depart from the mean by more than $\pm 1\%$.

Solar variability and forecasting. Smithsonian Misc. Coll., vol. 77, No. 5, pp. 16-18.
 Brenet, David. The combination of observations. (Cambridge, 1917) p. 33.

of the mathematical development and practical applica-

 $^{^2}$ Jørgensen, N. R., Undersøgelser over Frequensflader og Korrelation, Copenhagen, 1916.

than five observations were neglected, it appears that there can be no serious objection to such a procedure.

Table 2 shows, by months, the probability of a. m. temperature changes exceeding the stationary limits used in the verification of forecasts. The small chance of a temperature change exceeding 6° in the summer months is noteworthy. In Augustiboth the observed frequency

and the calculated frequency of such changes amount to only 8 in 1,000.

Figure 1 shows the observed distribution and the Gram-Charlier curves of best fit, as well as the type equation, and all the constants. An interesting feature is the persistent negative skewness, even in those months when the annual march of temperature is downward. This nega-

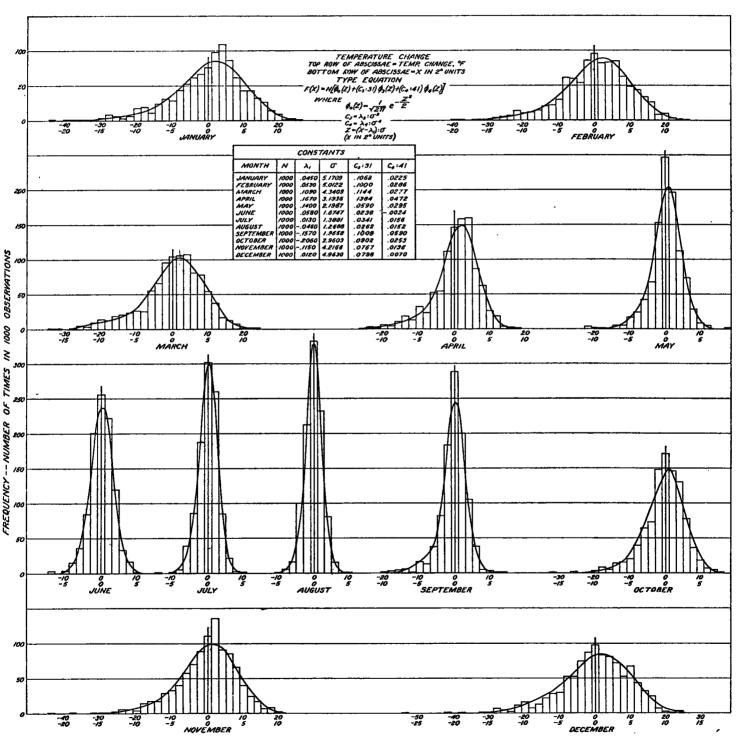


Fig. 1

tive skewness seems to be due, at least in part, to the steep temperature gradient in advance of passing Highs, giving rise to sudden and sharp falls in temperature, after which the rile is more gradual.

Table 1.— (A) Computation of parameters, January

Temper- ature change	F(X)	XF(X)	χı	X'F(X)	ιχ	X'F(X)	Ř	X'F(X)	(X+1)	(X+1)1 F(X)
-424343333432322824221810141210822202110810111011101110111011101110111011101110111011101110	20 1 19 1	0 - 200 - 200 - 200 - 200 - 200 - 200 - 200 - 201 - 20	324 2899 2256 2256 1968 169 121 1100 81 64 49 46 16 25 36 49 49 16 25 36 49 11 11 11 11 11 11 11 11 11 11 11 11 11	0 400 361 324 256 900 1,764 1,183 432 1,539 704 1,176 1,116 725 704 477 248 657 0 988 441	-10, 648 -9, 261 -8, 000 -6, 859 -5, 832 -4, 913 -4, 096 -3, 375 -2, 744 -1, 728 -1, 131 -1, 000 -125 -343 -216 -125 -64 -27 -8 -11 -1 -8 -27 -8 -125 -216 -343 -216 -343 -216 -343 -216 -343 -313 -1, 729 -313 -1, 729 -313 -1, 729 -313 -313 -313 -313 -313 -313 -313 -31	0	194, 481 160, 000 130, 321 104, 976 83, 521 65, 536 55, 625 38, 416 28, 561 10, 000 6, 561 11, 296 2, 401 11, 296 11,	0 180, 302 1 104, 976 83, 521 104, 976 83, 521 104, 976 83, 521 202, 500 124, 659 157, 624 40, 176 14, 293 67 7, 647 15, 616 26, 875 75, 782 106, 486 107, 487 11, 264 17, 7047 15, 616 100, 900 114, 641 41, 474 41, 474 41, 474 41, 474 41, 474 176 180, 322 106, 486 100, 900 114, 641 41, 474 41, 474 174 174 174 174 174 174 174 174 174	160, 000 130, 321 104, 976 83, 521 105, 536 55, 625 38, 416 28, 561 20, 736 14, 641 10, 000 11, 296 625 12, 401 11, 296 625 11, 404 11, 296 625 11, 404 11, 296 625 11, 404 11, 296 625 11, 404 11, 296 625 11, 404 11, 296 625 11, 296 625 11, 296 625 11, 296 625 12, 401 14, 641 10, 000 14, 641 10, 000 14, 641 10, 073 14, 641 10, 073 14, 641 15, 736 14, 641 16, 736 17, 736 18, 416	130, 321 104, 116 83, 621 65, 582 55, 625 153, 664 43, 923 77, 624 26, 411 19, 375 7, 424 3, 504 848 8, 910 22, 272 38, 125 55, 728 100, 243 131, 072
$\begin{array}{cccccccccccccccccccccccccccccccccccc$										

Table 1.—(B) Solution, January

x	X -λ ₁	$\frac{(x-\lambda_1)}{\sigma}$	¢0 (Z)	φ ₈ (Z)	φ4 (Z)	7	8	9	10	11
-22	23. 045 22. 045	-4.263	0. 0000 . 0001	0. 0018 . 0048	0. 0062 . 0150	0. 0002 . 0005	0. 0001 . 0003	0. 0003 . 0009	0	0
-20 -19	-21.045 -20.045 -19.045	-3. 877 -3. 683	. 0001 . 0002 . 0005	. 0069 . 0100 . 0177	. 0205 . 0300 . 0481	. 0007 . 0011 . 0019	. 0005 . 0007 . 0011	. 0013 . 0020 . 0035	0 0 1	1
$-17 \\ -16$	-18.045 -17.045 -16.045 -15.045	-3. 296 -3. 103	. 0009 . 0017 . 0032 . 0058	. 0290 . 0452 . 0665 . 0920	. 0707 . 0980 . 1228 . 1382	. 0031 . 0048 . 0071 . 0098	. 0016 . 0022 . 0028; . 0031	. 0056 . 0087 . 0131	1 2 3	1
$-14 \\ -13$	-15.045 -14.045 -13.045 -12.045	-2.716 -2.523	. 0100 . 0166 . 0264	. 1186 . 1404 . 1496	. 1312 . 0882 0027	. 0126 . 0149 . 0159	. 0030 . 0020 0001	. 0256 . 0335 . 0422	5 5 8	9
-11	-11. 045 -10. 045	-2.123 -1.943	. 0419 . 0605 . 0864	. 1335 . 0909 —. 0087	1562 3261 5272	. 0142 . 0097 —. 0009	0035 0073 0119	. 0526 . 0629 . 0736	10 12 14	7 18
-8 -7 -6	-7. 045 -6. 048	-1.556 -1.363 -1.169	. 1190 . 1575 . 2015	1071 2453 3933	6736 7398 6622	0114 0261 0418	0152 0166 0149	. 0924 . 1148 . 1448	18 22 28	11 24 31
-5 -4 -3	-4. 045 -3. 045	782 589	. 3354	4965 5488 5240	4631 0864 . 3661	0528 0583 0557	0104 0019 . 0082	. 1846 . 2337 . 2879		44 53
-2 -1 0	-1.045 045	202 009	. 3909	4144 2335 0107	. 7896 1. 0776 1. 1965 1. 0963	0441 0248 0011 0228	. 0178 . 0242 . 0262 . 0247	. 3427 . 3903 . 4240 . 4397	66 76 82 85	67 77
2 3 4	1. 955 2. 955	. 378		. 3987 . 5175 . 5499	. 8088 . 3901 —. 0503	. 0424 . 0550 . 0585	. 0182 . 0088 0011	. 4320 . 4027 . 3552	83 78 69	110 87
5 6 7 8	4. 955	. 958 1. 152 - 1. 345	. 2511 . 2054 . 1614	. 5029 . 3959 . 2586	4195 6557 7397	. 0535 . 0421 . 0275	0094 0148 0166	. 2952 . 2327 . 1723	57 45 33	43 43 32
9 10	8, 955 9, 955	1. 732 1. 925	. 0890 . 0625	. 1193 . 0000 —. 0849	6843 5341 3441	. 0127) . 0000 —. 0090	0154 0120 0077	. 1195 . 0770 . 0458	23 15 9	15 10
11 12 13 14	11, 955 12, 955	2.319 2.513	. 0416 . 0271 . 0169 . 0102	1344 1494 1413 1201	1536 0092 . 0945 . 1299	0143 0159 0151 0128	0035 0002 . 0019 . 0029	. 0238 . 0110 . 0037 . 0003	5 2 1	
Σ.	10. 900	2. 700	. 0102	1201	. 1289	0128	.0029	5. 1706	1,000	

Table 2.—The probability of temperature changes (7 a. m. to 7 a. m.) at Montgomery, Ala.

	Sta- tion- ary limit		Excee	Within station- ary limit					
Month		R	ise	Fall		Combined			
	111111	Actual	Theo- retical	Actual	Theo- retical	Actual	Theo- retical	Actual	Theo- retical
January. February	10	13. 0 12. 8	13. 3 12. 2	13. 9 12. 1	13. 4 12. 5	26. 9 24. 9	26. 7 24. 7	73. 1 75. 1	73. 8 75. 3
March April	10	7. 9 5. 1	8.9 5.6	11.3 9.2	10. 3 8. 6	19. 2 14. 3	19. 2 14. 2	80. 8 85. 7	80. 8 85. 8
May June July	6	1.5 2.0	1.6 1.5	3. 2 2. 4	3.0 2.0 1.0	4.7 4.4 1.2	4.6 3.5 1.4	95. 3 95. 6 98. 8	95.4 96.4 98.4
August September October	ñ	1.2	1.6	4.9 7.5	.6 6.0	.8 6.1	.8 7.6	99, 2 93, 9	99. : 92.
October November December		3. 3 10. 6 12. 1	3. 8 12. 7 12. 8	7. 5 14. 1 12. 3	8. 2 14. 4 13. 4	10.8 24.7 24.4	12. 0 27. 1 26. 2	89. 2 75. 3 75. 6	88. 72. 73.
Sums Means		70. 1 5. 8	74. 6 6. 2	92.3	93. 4 7. 8	162, 4 13, 5		1, 037. 6 86. 5	1, 032. 86.